

Computation of MHD Parameters in the LFM model

1. Pressure

In ideal MHD model [Gombosi, 1998],

$$\varepsilon = \frac{1}{2} \rho v^2 + \frac{p_{th}}{\gamma - 1} \quad (1)$$

where ε is the sum of flow and thermal energy density,

ρ is mass density,

v is flow velocity,

p_{th} is thermal pressure, and

$\gamma = 5/3$, is the polytropic index.

Eq. (1) can be rearranged as,

$$\begin{aligned} \frac{2}{3} \varepsilon &= \frac{1}{3} \rho v^2 + p_{th} \\ &= p_{dyn} + p_{th} \end{aligned} \quad (2)$$

where $p_{dyn} = \text{dynamic pressure} = \frac{1}{3} \rho v^2$

$p_{th} = \text{thermal pressure} = nkT$

2. Mean Energy

Mean energy, \bar{E} , is given by,

$$\bar{E} = \frac{\varepsilon}{n} \quad (3)$$

where n is number density. Combining Eqs. (2) and (3), \bar{E} can be obtained by,

$$\bar{E} = \frac{3}{2} \frac{p_T}{n} \quad (4)$$

where $p_T = p_{dyn} + p_{th}$ is the total pressure.

Reference:

Gombosi, T. I, *Physics of the Space Environment*, Cambridge, New York, 1998